# **SURE-based Blind Gaussian Deconvolution** Feng Xue and Thierry Blu

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## Summary

- **Problem**: blind deconvolution without the knowledge of the Point Spread Function;
- **Basic procedure**: PSF estimation + non-blind deconvolution with estimated PSF;
- Our scope: Gaussian PSF with unknown variance  $s_0^2$  (to be estimated);
- **Originality**: novel objective functional blur SURE, a modified version of SURE (Stein's unbiased risk estimate);
- **Potential**: possibly extend SURE-based framework to other types of PSF with known parametric form.

## **Problem statement**

### Linear observation model

 $\mathbf{y} = \mathbf{H}_0 \mathbf{x} + \mathbf{n}$ 

where

- $\mathbf{H}_0$  the latent true convolution matrix associated with true PSF  $\mathbf{h}_0$
- Gaussian noise  $\mathbf{n} \sim \mathcal{N}(0, \sigma^2 \mathbf{I})$



**Problem**:  $\mathbf{x} = ?$  and  $\mathbf{h}_0 = ?$ , knowing  $\mathbf{y}$  only. Solution — separate estimation of PSF, and then signal: Step 1 — PSF estimation; Step 2 — deconvolution<sup>\*</sup>.  $\star$  We use our recently proposed SURE-LET approach to perform (non-blind) deconvolution [1,2].

### Gaussian kernel

• Parametric form with standard deviation s

$$\mathbf{h}(i, j; \mathbf{s}) = C \cdot \exp\left(-\frac{i^2 + j^2}{2\mathbf{s}^2}\right)$$

$$\mathbf{s} - \text{blur size, width of the Gaussian shape;}$$

$$C - \text{normalization coefficient, s.t. } \sum_{i,j} \mathbf{h}(i, j) = 1.$$

- $\mathbf{h}_0$  latent true Gaussian kernel with unknown width  $s_0$
- Question: how to estimate  $s_0$ , from observed y?

[1]. F. Xue, F. Luisier, and T. Blu, SURE-LET image deconvolution using multiple Wiener filters, *ICIP 2012*. [2]. F. Xue, F. Luisier, and T. Blu, Multi-Wiener SURE-LET Deconvolution, *submitted to IEEE TIP*.

## Blur SURE as a new criterion

• blur MSE (mean squared error) is defined as (with unknown  $H_0x$ ):

blur MSE =  $\frac{1}{N} \mathscr{E} \left\{ \| \mathbf{HF}(\mathbf{y}) \right\}$ 

• blur SURE — unbiased estimate of the blur MSE:

### **Remarks**:

-the blur SURE depends on the observed data only (NOT on  $\mathbf{H}_0$  and  $\mathbf{x}$ ); -divergence operator:  $\operatorname{div}_{\mathbf{y}}\mathbf{u} = \sum_{n=1}^{N} \frac{\partial u_n}{\partial u_n}$  for  $\forall \mathbf{u} \in \mathbb{R}^N$ ; - Minimizing the blur-SURE yields results that are very close to minimizing the blur-MSE.

## **Blur-SURE** minimization for Wiener processing

**Theorem**: Consider the approximate Wiener filtering:

$$\mathbf{Y}(\mathbf{y}) = (\mathbf{H}^{\mathrm{T}}\mathbf{H} + \lambda \mathbf{I})^{\mathrm{T}}$$

 $\mathbf{W}_{\mathbf{H},\lambda}$ Then, the minimization of the blur MSE over both H and  $\lambda$ : blur MSE

$$\min_{\mathbf{H},\lambda} \frac{1}{N} \| \mathbf{H} \mathbf{W}_{\mathbf{H},\lambda} \mathbf{y} - \mathbf{w} \|$$

yields  $\mathbf{H} pprox \mathbf{H}_0$ .

### Explanation (Fourier representation)

Consider the exact Wiener processing with known  $H_0(\omega)$ :

$$W(\omega) = rac{H_0^*(\omega)}{|H_0(\omega)|^2 + \sigma^2/\omega}$$

where  $S(\omega)$  is the power spectrum density of image x. Then,  $U_0(\omega) = H_0(\omega)W(\omega)$  behaves like a band indicator.

The blur-SURE minimization results in another band indicator  $U = HW_{H,\lambda}$ , which is as close as possible to  $\mathbf{U}_0$ :



### Approximation of the band indicator $U_0(\omega)$





$$)-\mathbf{H}_{0}\mathbf{x}\left\Vert ^{2}
ight
angle$$

 $\epsilon = \frac{1}{N} \left\| \mathbf{HF}(\mathbf{y}) - \mathbf{y} \right\|^2 + \frac{2\sigma^2}{N} \operatorname{div}_{\mathbf{y}} \left( \mathbf{HF}(\mathbf{y}) \right) - \sigma^2$ 

 $|\mathbf{H}_0\mathbf{x}|$ 





-150 -100 -50 0 50 100 150 A typical example for  $U_0(\omega)$ 



## **Results and discussions**

### SURE-based framework to estimate $s_0$ and $\lambda$ $\mathbf{W}_{\mathbf{H}}$ tentative $\stackrel{s}{,} \longrightarrow$

<u> </u>	/	$\mathbf{\Lambda}$	$\cdots$ <b>II</b> , $\wedge$
			Wiener filterir
			Stage 1: PSF es
~			<b>-</b>

 $\star$  One possibility is to use alternating minimizations between s and  $\lambda$ .

### Estimation of $s_0$ , followed by deconvolution

Table 1: Blind deconvolution (Cameraman)												
BSNR (in dB)	40	30	20	10	40	30	20	10	40	30	20	10
true $s_0$	$s_0 = 1.0$			$s_0 = 2.0$				$s_0 = 3.0$				
estimated $s_0$	1.12	1.19	1.24	1.33	2.15	2.18	2.25	2.48	3.28	3.34	3.37	3.52
PSNR difference•	0.26	0.18	0.12	0.09	0.11	0.07	0.07	0.10	0.13	0.11	0.08	0.10
PSNR difference after deconvolution with eracle												

PSINK difference after deconvolution with oracle

ble 2: S	SNR improvement (in dB) of deconvolution performance for										
	Method	SAR1 [3] SAR2 [3] TV1 [4] TV2 [4] SURI									
	BSNR	40dB									
	Cameraman	1.03	1.01	1.82	1.73	3.15					
	Lena	1.35	1.43	2.53	2.59	4.54					
	BSNR			20dB							
	Cameraman	1.16	-8.83	1.70	-40.89	2.15					
	Lena	1.62	-11.32	2.62	-32.50	3.13					

### A visual example

blurred with  $s_0 = 1.50$ BSNR = 20dBPSNR = 23.23dB



### Real data

Restoration of *Jupiter*: • The estimated noise std is  $\sigma = 4.68$  by using MAD (median absolute deviation); • Estimated  $s_0 = 2.41$ 



[3]. R. Molina, J. Mateors, and A. Katsaggelos, *IEEE TIP*, vol.15, no.12, pp.3715–3727, 2006. [4]. S. Babacan, R. Molina, and A. Katsaggelos, *IEEE TIP*, vol.18, no.1, pp.12–26, 2009

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Note that the PSNR loss due to the inexactness of the estimation is kept within 0.2dB. Comparisons with the state-of-the-art in blind deconvolution

$s_0^2 = 9$	nce for	rma	perfo	b) of deconvolution p				dB	in
	SURE	[4]	TV2	[4]	TV1	[3]	SAR2	[3]	1

### blind deconvolution with estimated $s_0 = 1.79$ PSNR = 25.65 dB



non-blind deconvolution with known  $s_0 = 1.50$ PSNR = 25.75dB



Observed image



